# Rotational spectroscopy of *n*-propanol: Aa and Ag conformers\*

O. Zingsheim<sup>1</sup>, J. Maßen<sup>1</sup>, H. S. P. Müller<sup>1</sup>, B. Heyne<sup>1</sup>, M. Fatima<sup>1</sup>, L. Bonah<sup>1</sup>, A. Belloche<sup>2</sup>, F. Lewen<sup>1</sup>, and S. Schlemmer<sup>1</sup>,

<sup>1</sup> I. Physikalisches Institut, Universität zu Köln, Zülpicher Straße 77, 50937 Köln, Germany

e-mail: zingsheim@ph1.uni-koeln.de; hspm@ph1.uni-koeln.de

<sup>2</sup> Max-Planck-Institut für Radioastronomie, Auf dem Hügel 69, 53121 Bonn, Germany

Received 2022, A&A accepted doi:10.1051/0004-6361/202243571

#### ABSTRACT

Context. The primary alcohol n-propanol (i.e., normal-propanol or propan-1-ol; C<sub>3</sub>H<sub>7</sub>OH) occurs in five different conformers: Ga,  $G_{g}, G_{g'}, A_{a}$ , and  $A_{g}$ . All rotational spectra of the three conformers of the G family are well described, making astronomical search of their spectroscopic signatures possible, as opposed to those of the Aa and Ag conformers.

Aims. Our goal is to facilitate the astronomical detection of Aa and Ag conformers of n-propanol by characterizing their rotational

Methods. We recorded the rotational spectra of n-propanol in the frequency domain of 18-505 GHz. Additional double-modulation double-resonance (DM-DR) measurements were performed, more specifically with the goal to unambiguously assign weak transitions of the Aa conformer and to verify assignments of the Ag conformer.

*Results.* We derived a spectroscopic quantum mechanical model with experimental accuracy (with  $J_{\text{max}} = 70$  and  $K_{a,\text{max}} = 6$ ) for Aa*n*-propanol. Furthermore, we unambiguously assigned transitions (with  $J_{\text{max}} = 69$  and  $K_{a,\text{max}} = 9$ ) of Ag *n*-propanol; in doing so, we

Conclusions. The astronomical search of all five conformers of n-propanol is now possible via their rotational signatures. These are applied in a companion article on the detection of *n*-propanol toward the hot molecular core Sgr B2(N2).

Key words. molecular data – methods: laboratory: molecular – techniques: spectroscopic – ISM: molecules – radio lines: ISM –

Received 2022, A&A accepted doi:10.1051/0004-6361/202243571 **ABST Context.** The primary alcohol *n*-propanol (i.e., normal-propanol of *Gg, Gg', Aa*, and *Ag*. All rotational spectra of the three conformer of their spectroscopic signatures possible, as opposed to those of th *Aims.* Our goal is to facilitate the astronomical detection of *Aa* as spectra. *Methods.* We recorded the rotational spectra of *n*-propanol in the double-resonance (DM-DR) measurements were performed, more so of the *Aa* conformer and to verify assignments of the *Ag* conformer *Results.* We derived a spectroscopic quantum mechanical model v *n*-propanol. Furthermore, we unambiguously assigned transitions of prove the existence of two tunneling states, *Ag\** and *Ag<sup>-</sup>. Conclusions.* The astronomical search of all five conformers of *n*-applied in a companion article on the detection of *n*-propanol towa **Key words.** molecular data – methods: laboratory: molecular – submillimeter: ISM **1. Introduction** To date, more than 260 molecules have been detected in space.<sup>1</sup> Many rotational molecular spectra are collected in databases such as the Cologne Database for Molecular Spectroscopy (CDMS) (Müller et al. 2001, 2005; Endres et al. 2016), allow-ing astronomers to search for their spectroscopic signatures in space. Several complex organic molecules (COMs) have been found thus far, for instance, in the warm gas around proto-stars in star forming regions, such as formamide (Rubin et al. 1971), ethyl cyanide (Johnson et al. 1977), and ethanediol (Hol-lis et al. 2002). One prominent example is the Protostellar In-terferometric Line Survey (PILS; Jørgensen et al. 2016) with the Atacama Large Millimeter/submillimeter Array (ALMA). Among the COMs, alcohols have also been found in space. So far, methanol (CH<sub>3</sub>OH; Ball et al. 1970) and ethanol (C<sub>2</sub>H<sub>5</sub>OH; Zuckerman et al. 1975; Pearson et al. 1997) have been detected. Ethanol occurs as two conformers, so-called anti-ethanol (Zuck-erman et al. 1975) and ga Ethanol occurs as two conformers, so-called anti-ethanol (Zuckerman et al. 1975) and gauche-ethanol (Pearson et al. 1997), and both have been referenced based on their first detection. The next heavier alcohol, propanol (C<sub>3</sub>H<sub>7</sub>OH), is thus another candidate for a detection in space, particularly in environments with large amounts of ethanol. The possible detection of propanol may be further supported by the detection of its isomer, ethyl methyl

ether (or methoxyethane, CH<sub>3</sub>CH<sub>2</sub>OCH<sub>3</sub>) in space (Tercero et al. 2015, 2018; Manigand et al. 2020).

Propanol itself occurs in two isomers, as a primary alcohol *n*-propanol (normal-propanol, propan-1-ol, normal propyl alcohol; CH<sub>3</sub>CH<sub>2</sub>CH<sub>2</sub>OH), where the hydroxyl group (-OH) is attached to the outer carbon atom, and as a secondary alcohol *i*-propanol (iso-propanol, propan-2-ol, or iso-propyl alcohol; CH<sub>3</sub>CH(OH)CH<sub>3</sub>), where the hydroxyl group is attached to the central carbon atom. n-Propanol, the subject of this study, occurs in five different conformers: Ga, Gg, Gg', Aa, and Ag, see Abdurakhmanov et al. (1968, 1970, 1976); Dreizler & Scappini (1981); Lotta et al. (1984); Maeda et al. (2006); Kisiel et al. (2010) and references therein. The capital letters G and A refer to the rotation of the heavy nuclei plane C-C-C compared to C-C-O. An "anti" (A) configuration<sup>2</sup> describes a rotation by  $180^{\circ}$  and "gauche" (G) by about  $60^{\circ}$ . The small letters refer to the rotation of the dihedral angle of the hydroxy group (-OH).

The G family of conformers (Ga, Gg, and Gg') of n-propanol has been rather extensively studied in rotational spectroscopy, as in, for instance, Abdurakhmanov et al. (1970); Maeda et al. (2006); Kisiel et al. (2010) and references therein. A single state analysis of Ga was able to fit 2865 lines up to 375 GHz, but systematic deviations occurred for some transitions with higher Jquantum numbers (Maeda et al. 2006). Based on this observation, a combined analysis of the three G conformers was carried

<sup>\*</sup> The supplementary material, our predictions of the rotational spectra (\*.cat files), are only available at CDS.

<sup>&</sup>lt;sup>1</sup> https://cdms.astro.uni-koeln.de/classic/molecules

 $<sup>^2</sup>$  The term "trans" (T) was often used instead of "anti" historically, although it should only be used to refer to planar molecules.

out by taking into account Coriolis interaction and Fermi resonances (Kisiel et al. 2010). Thanks to the quantum mechanical models of rotational spectra with experimental accuracy (Kisiel et al. 2010), conformers of the *G* family can be found in space. Furthermore, this distinctive analysis gives confidence that the *G* and *A* conformers can be treated separately as within the *G* family, the conformers seem to be unaffected by conformers of the *A* family (*Aa* and *Ag*).

The ab initio calculations have shown some scatter for the relative energies of *n*-propanol conformers, as, for instance, in Lotta et al. (1984); Kahn & Bruice (2005) and references therein (a detailed overview is given in Maeda et al. (2006), but they agree in the fact that all conformers, in particular from both G and A families, are rather close in energy).<sup>3</sup> Considering only Boltzmann statistics, the Aa and Ag conformers are expected to have about the same chance of being detected in the warm interstellar medium as the G family conformers.

The most advanced rotational analysis of the Aa n-propanol was derived by using a Fourier transform microwave (FTMW) spectrometer (8-18 GHz) and a Stark modulated microwave spectrometer (up to 40 GHz) (Dreizler & Scappini 1981). Four conformers of *n*-propanol occur as mirror configurations (twofold degeneracy due to "left-" and "right-handed" versions,  $C_1$ symmetry), except for the Aa conformer, which has a plane of symmetry and one-fold degeneracy ( $C_s$  symmetry), as shown in Fig. 1. The Ag conformer occurs in mirror images (see Fig. 1), however, the barrier for hydrogen tunneling between the two equivalent Ag forms is low enough to distinguish symmetric  $Ag^+$ and antisymmetric  $Ag^-$  tunneling states (Abdurakhmanov et al. 1976; Kisiel et al. 2010). The tunneling transition state has  $C_s$ symmetry and tunneling-rotation interactions need to be taken into account, as has been done, for instance, for gauche-ethanol (Pearson et al. 2008) and gauche-propanal more recently (Pickett & Scroggin 1974; Zingsheim et al. 2017, 2022). So far, only eleven doublet features up to 30 GHz were tentatively assigned to  $Ag^+$  and  $Ag^-$  tunneling states with  $J_{\text{max}} = 4$  (Abdurakhmanov et al. 1976).

The main aim of the present study is to facilitate the astronomical detection of Aa *n*-propanol and Ag *n*-propanol. For this purpose, the rotational spectra are measured in selected frequency ranges from 18 to 505 GHz. The employed experimental setups are presented in Sect. 2. The experimental data are analyzed in Sect. 3 with a focus on secure assignments of the perturbed systems. Next, the spectroscopic results and the astronomical search of *n*-propanol are discussed in Sect. 4. Finally, our main conclusions are given in Sect. 5.

# 2. Experimental techniques

We recorded rotational spectra of *n*-propanol (i) in the 17.5–26.5 GHz region using chirped-pulse Fourier transform microwave (CP-FTMW) jet spectroscopy and (ii) in the regions of 33–67 GHz, 70–129 GHz, and 160–505 GHz using conventional 2f demodulation absorption techniques. Additional (iii) double-modulation double-resonance (DM-DR) measurements support assignments of weaker transitions to the *Aa* conformer and, in particular, secure the identification of various  ${}^{q}R$  se-

Fig. 1: Aa and Ag conformers of *n*-propanol. Hydrogen, carbon, and oxygen atoms are depicted with white, black, and red spheres, respectively. The atoms C-C-C-O-H span a (symmetry) plane for the Aa conformer ( $C_s$  symmetry) and it is shown as face-on and side view for clarification. Then, Ag *n*-propanol occurs in two mirror-image forms leading to the  $Ag^+$  and  $Ag^-$  tunneling states. The *G* conformers are not depicted here, but the oxygen is out of the heavy nuclei plane for all of them.

ries<sup>4</sup> of the *Ag* conformer. The (i) CP-FTMW, (ii) conventional (sub-)millimeter wave (MMW) absorption spectrometers, and (iii) the DM-DR setup are introduced in detail in (i) Hermanns et al. (2019), (ii) Ordu et al. (2019); Martin-Drumel et al. (2015), and (iii) Zingsheim et al. (2021), respectively. Thus, their main characteristics are only briefly summarized in the following.

In terms of (i) the 17.5-26.5 GHz range was covered in steps of 1 GHz bandwidth using a CP-FTMW spectrometer. In contrast to the above-described setup (Hermanns et al. 2019), mixing up the CP and mixing down the molecular signal is no longer required as the CP is generated directly in the target frequency region using an arbitrary waveform generator (Keysight M8195A 65 GSa/s). In the present setup, each 1 GHz broad CP spanned 2 µs and was amplified with a 4 W amplifier. These amplified pulses were then broadcast via a horn antenna to excite the molecular ensemble. After the pulse was stopped, the resulting free induction decay (FID) of the molecules was collected with another horn antenna, amplified using a low noise amplifier, and recorded for 10 µs on a 100 GS/s oscilloscope. Each measurement consists of 65 000-100 000 acquisitions, which were performed with a 10 Hz repetition rate. The averaged FIDs were then Fourier transformed using a Hanning window, which resulted in a frequency resolution of about 290 kHz, considering the full width at half maximum (FWHM). The sample was prepared by mixing 1 mL of liquid *n*-propanol (Sigma-Aldrich, 97.7 % purity) with 7 bar neon in a 10 L container. The mixture was supersonically expanded with a relative stagnation pressure of 1 bar and an opening duration of 500 µs.

For (ii) either an Agilent E8257D or a Rohde & Schwarz SMF100A synthesizer generated electromagnetic waves. Electromagnetic waves were directly coupled to a horn antenna to record single spectra in the 33–67 GHz region. We used in-house developed electronics for operating an amplifier tripler chain in full saturation mode to fully cover the 70–129 GHz region and used commercially available cascaded doublers and triplers from Virginia Diodes Inc. (VDI) to fully cover the 160–505 GHz region. Frequency modulation (FM) increases the signal-to-noise ratio (S/N) and the 2f demodulation creates absorption features

<sup>&</sup>lt;sup>3</sup> For instance, the *Ga* conformer was predicted to be the energetically lowest energy conformer with the others higher in energy by E(Aa) = $24 \text{ cm}^{-1}$  (35 K),  $E(Gg) = 35 \text{ cm}^{-1}$  (50 K),  $E(Ag) = 42 \text{ cm}^{-1}$  (60 K), and  $E(Gg') = 49 \text{ cm}^{-1}$  (70 K) (Kahn & Bruice 2005). Experimentally derived values for the *G* family are  $E(Ga) = 0 \text{ cm}^{-1}$ ,  $E(Gg) = 47.8 \text{ cm}^{-1}$ (69 K), and  $E(Gg') = 50.8 \text{ cm}^{-1}$  (73 K) (Kisiel et al. 2010).

<sup>&</sup>lt;sup>4</sup> Transitions with identical selection rules are called a series, here abbreviated as  ${}^{\Delta K_a}\Delta J$ , with  $\Delta J$  represented by capital letters *P*, *Q*, and *R* for  $\Delta J = -1, \pm 0$  or +1, respectively, and  $\Delta K_a = -1, \pm 0$ , or +1 by superscripts *p*, *q*, or *r*, respectively. For instance, *a*-type transitions ( $\Delta K_a = 0$ ) with  $\Delta J = +1$  are called  ${}^{q}R$  series transitions. A subscript may be used additionally to highlight the respective  $K_a$  quantum number ( ${}^{q}R_{K_a}$  series).

that are close to the second derivative of a Voigt profile. We used a modulation frequency of  $f \approx 47$  kHz and its amplitude is on the order of the FWHM of the absorption profiles. The absorption path was 14 m in a single pass configuration for measurements up to 130 GHz (Ordu et al. 2019) and 10 m in double pass configuration otherwise (Martin-Drumel et al. 2015). Here, *n*-Propanol was vaporized into the cell to pressures in the range of  $10-40 \,\mu$ bar.

For (iii) we used both synthesizers, namely, Agilent as the so-called probe and Rohde & Schwarz as the pump, simultaneously for DM-DR measurements; a modulated pump radiation was applied in addition to the frequency modulated probe radiation. The probe signal was usually frequency tripled, and inhouse developed electronics used for operation in full saturation mode ensured a stable output power of about 1 mW (0 dBm). The pump source, an active frequency multiplier (AFM6 70-110+14 from RPG Radiometer Physics), delivered output powers of up to 60 mW (17.8 dBm). The final pump and probe frequencies were both located in the W-band. We performed some additional measurements with the probe frequency being around 200 GHz using commercial components by VDI. It is only the probe transitions that share an energy level with the pump transition that are affected by the additional source. An additional 1f' demodulation of the pumped signal is experimentally realizing a difference spectrum of the on- and off-resonant measurements, or, in other words, a subtraction of a conventional and a DR measurement. In this way, it is only the lines of interest remaining and the assignment of even weak or blended features is facilitated (Zingsheim et al. 2021).

# 3. Results: Spectroscopic signatures

The measurement results for Aa and Ag *n*-propanol are presented in this section. The central focus of this chapter is the analyses and the resulting quantum mechanical models of rotational spectra, as well as the verification of assignments, since both conformers show perturbed patterns. Our assignment procedure benefited, in terms of speed and unambiguity, from the *LLWP* software, which is based on Loomis-Wood plots<sup>5</sup> (Bonah et al. 2022). Adjacent transitions of certain series are plotted around their predicted frequency in Loomis-Wood plots to increase the confidence of correct assignments or to visualize systematic deviations between observed and predicted transition frequencies.

# 3.1. Aa n-propanol

The Aa conformer is an asymmetric rotor quite close to the prolate limit with  $\kappa = -0.978 (\kappa = [2B - A - C]/[A - C])$ . The dipole moments were determined with Stark effect measurements to be  $\mu_a = 0.21(7)$  D and  $\mu_b = 1.54(2)$  D (Abdurakhmanov et al. 1970). In total, 73 lines (118 transitions) were assigned and fit up to 40 GHz in a previous study (Dreizler & Scappini 1981). Therein, all lines from Abdurakhmanov et al. (1968) were remeasured; a- and b-type transitions were assigned and fit and in particular the assignments of weak *a*-type transitions were secured with the help of MW-MW DR measurements (Dreizler & Scappini 1981). A splitting due to the internal rotation of the methyl group (-CH<sub>3</sub>) could sometimes be resolved (14 A+14 E transitions) and together with 45 blended doublet features, the transition frequencies were reproduced with a standard deviation of 195 kHz (Dreizler & Scappini 1981). Thereby, the potential barrier of the methyl group internal rotation was determined



Fig. 2: Loomis-Wood plot of *Aa n*-propanol - Screenshot of the *LLWP* software (Bonah et al. 2022). Shown here is a doublet series (A and E components of the methyl group internal rotation) of the  $J_{K_a,K_c} = J_{3,J-3} \leftarrow J_{2,J-2}$  *b*-type transitions from J = 6 (bottom panel) to J = 36 (top panel). Predictions of the A components are set as center frequencies and are depicted by green sticks. The E components are depicted by red ones. This <sup>*r*</sup>*Q* series with  $K_a + K_c = J$  and  $\Delta K_a, \Delta K_c = +1, -1$  is well reproduced by our spectroscopic model described in Table 2 up to J = 22, however, transitions up to J = 41 ( $\Delta \nu \approx 90$  MHz) are straightforwardly assignable.

to be  $V_3 = 955(21) \text{ cm}^{-1}$  (or 2730(60) cal/mol) and the angle of the internal rotation axis *i* to the principal moment of inertia axis *a* to be  $\langle (i, a) = (29 \pm 1)^\circ$  (Dreizler & Scappini 1981).

We used the available parameters and transition frequencies (Dreizler & Scappini 1981) to set up an initial spectroscopic model of an asymmetric rotor, including three-fold internal rotation in ERHAM (Groner 1992, 1997, 2012). Thanks to the initial prediction and with the application of an iterative assignment and fit strategy, we could assign strong and also weak b-type transitions in a straightforward fashion (see Fig. 2). Transitions from Dreizler & Scappini (1981) were remeasured if possible.<sup>6</sup> Furthermore, we secured assignments of weak a-type transitions in the W-band region with the DM-DR technique. Additionally, transitions of the <sup>r</sup>R series  $(J_{K_a,K_c} = J_{4,J-3} \leftarrow (J-1)_{3,J-4}$  with J up to 13) have been unambiguously assigned above 200 GHz with DM-DR measurements (see Fig. 3). The assignment of internal rotation components, A and E, is unambiguous as we observed E\* transitions, which are forbidden in rigid rotors (Herschbach & Swalen 1958). However, the  $^{r}R$  series transitions appeared to be noticeably shifted from their prediction and not all of them are incorporated in the final fit. Including these transi-

<sup>&</sup>lt;sup>5</sup> Download at https://llwp.astro.uni-koeln.de/

<sup>&</sup>lt;sup>6</sup> Two *Q*-branch *a*-type lines (or four transitions) were not included in the analysis of Dreizler & Scappini (1981), but were secured by them with MW-MW DR spectroscopy; These lines are successfully incorporated to the fit in this study. However, six other transitions have been unweighted in this study due to conspicuously large deviations.



Fig. 3: DM-DR measurements of *Aa n*-propanol of the  $J_{K_a,K_c} = 10_{4,6} \leftarrow 9_{3,7}$  transition around 232.55 GHz. The lower panels show two DM-DR measurements where the A and E components of the  $J_{K_a,K_c} = 9_{3,7} \leftarrow 9_{2,8}$  transition are used as pump frequencies at 113811.6420 MHz and 113810.5391 MHz, respectively. A conventional measurement is shown in the upper panel. E\* transitions, which are forbidden in rigid rotors (Herschbach & Swalen 1958), secure the assignment of internal rotation components. We note that slightly off-resonant pumping leads to small, asymmetric features of A and E components in the DM-DR spectrum of the other one, as pumped A and E components are rather close in frequency, cf. Zingsheim et al. (2021).

tions is possible if others are excluded instead. The final choice of fit transitions is somewhat ambiguous: the quantum number coverage is visualized in Fig. C.1 of the Appendix. Larger systematic deviations,  $\Delta v = v_{Obs} - v_{Calc}$ , occur from our final predictions for various series (cf. Fig. 2). These deviations start at similar J values for series involving identical  $K_a$  values, see Table 1. However, the assignment of these transitions is unambiguous as demonstrated in Fig. 2 and the assignment of internal rotation components is also secured (cf. Fig. 3). Finally, we could assign 1255 transitions with frequencies up to 505 GHz and  $J_{max} = 70$ , thereof, 928 transitions are well reproduced by the prediction. The determined spectroscopic parameters can be found in Table 2.

#### 3.2. Ag n-propanol

The three members of the *G* family were extensively described in the literature plus the *Aa* conformer to some extent. The situation is different for the fifth conformer of *n*-propanol: *Ag*. There is no spectroscopic model of rotational spectra with experimental accuracy available, and only 22  $J_{K_a,K_c} \leftarrow (J-1)_{K_a,K_c-1} {}^{q}R$ series transitions (1, 3, 5, and 2 transitions for J = 1, 2, 3, and 4, respectively, for each of the two tunneling states), were tentatively assigned in the MW region (Abdurakhmanov et al. 1976). Therein, doublet features were observed and assigned to the tunneling states  $Ag^+$  and  $Ag^-$ . The average B + C was determined to be about 7.3 GHz (Abdurakhmanov et al. 1976). A repeating signature with this spacing was also observed in the MMW region (see Fig. 2 of Kisiel et al. 2010) and can be attributed to  $K_a$  structures with increasing *J* quantum numbers of the *Ag* 

Table 1: Quantum number overview of the final analysis of Aa *n*-propanol.<sup>a</sup>

$K_a^{b}$	$K_a^{\prime\prime} + K_c^{\prime\prime}$	<i>a</i> -type <sup>c</sup>	<i>b</i> -type <sup>c</sup>			
		$^{q}R$	$^{r}P$	rQ	rR	${}^{p}R$
		$\pm 0, -1$	+1,-3	+1, -1	+1, -1	-1,+3
		$\pm 0, -1$	+1, -1	+1, -1	+1,+1	-1,+1
0	J	16 (16)	-	32 (32)	69 (70)	68 (70)
1	J + 1	17 (17)	7 (7)	28 (28)	20 (20)	29 (29)
1	J	16 (16)	7 (7)	31 (41)	31 (53)	41 (53)
2	J + 1	16 (16)	13 (13)	20 (26)	23 (32)	26 (35)
2	J	16 (16)	13 (13)	22 (41)	20 (37)	22 (26)
3	J + 1	16 (16)	-	9 (25)	9 (12)	-
3	J	16 (16)	-	9 (11)	9 (13)	-
					. ,	
4	J + 1	11 (11)	_	_	_	_
4	J	11 (11)	-	-	-	-

<sup>a</sup> Values represent  $J_{\text{max}}$ , the highest J quantum number of well reproduced transitions of a given series in our final spectroscopic model (Values in parentheses represent  $J_{\text{max}}$  of assigned transitions); <sup>*q*</sup>Q transitions are neglected here.

<sup>b</sup> The given  $K_a$  is always the lower one of  $K'_a$  and  $K''_a$ .

<sup>c</sup> Selection rules for  $K_a$  and  $K_c$  quantum numbers:  $\Delta K_a = K'_a - K''_a$ ,  $\Delta K_c = K'_c - K''_c$ . Upper values are for asymmetry side  $K''_a + K''_c = J'' + 1$  and lower ones for  $K''_a + K''_c = J''$ . Additionally,  $\Delta K_a$  with  $-1, \pm 0, +1$  is also represented by the superscripts p, q, r, respectively. The capital letters P, Q, R, represent  $\Delta J$  with  $-1, \pm 0, +1$  respectively.

conformer. However, to our knowledge no spectroscopic model with experimental accuracy nor assignments are publicly available for that latter work. Namely, *Ag n*-propanol is a perturbed system that cannot be analyzed easily. The focus of this work is on providing initial assignments of  ${}^{q}R$  series transitions, as the dipole moments are calculated to be  $\mu_a = 1.15 \text{ D}$ ,  $\mu_b = 0.39 \text{ D}$ , and  $\mu_c = 1.15 \text{ D}^7$ , but the energy difference of  $Ag^+$  and  $Ag^-$  tunneling states is unknown so far. Then, *Ag n*-propanol is expected to be close to the prolate limit with  $\kappa = -0.98$  (Kisiel et al. 2010).

We performed CP-FTMW spectroscopy in particular to check the assignments of  $3_{K_a,K_c} \leftarrow 2_{K_a,K_c-1}$  *a*-type transitions from Abdurakhmanov et al. (1976). We measured not only the aforementioned transitions of Ag *n*-propanol, but also of gauchepropanal as its resulting tunneling spectroscopic pattern is already well-understood (Zingsheim et al. 2022). Similar spectroscopic signatures are expected for both molecules and by comparing both spectra; literature assignments of Ag *n*-propanol could not be confirmed, but the first tentative assignments of typical doublet transitions, originating in  $Ag^+$  and  $Ag^-$  tunneling states, could be made (cf. Fig. B.1 in the appendix).

DM-DR measurements secured the linkages of various transitions that occur about every 7.3 GHz in the W-band region. Afterwards, we assigned  ${}^{q}R$  series transitions in the frequency region from 36 to 505 GHz with the help of the *LLWP* software (Bonah et al. 2022). Some of the  ${}^{q}R$  series of the Ag conformer are visualized exemplarily in a Fortrat diagram (Fig. 4). Assignments of  ${}^{q}R_{K_a}$  series transitions,  $J_{K_a,K_c} \leftarrow J - 1_{K_a,K_c-1}$ , are done as they start in rows of the Fortrat diagram with  $J = K_a + 1$  $(K''_a \leq J'')$ . Two patterns are found, with each consisting of a doublet series for each  ${}^{q}R_4$ ,  ${}^{q}R_5$ ,  ${}^{q}R_6$ ,  ${}^{q}R_7$ ,  ${}^{q}R_8$ , and  ${}^{q}R_9$  series

<sup>&</sup>lt;sup>7</sup> Allowed *a*- and *b*-type transitions occur within a tunneling state:  $Ag^+ \leftrightarrow Ag^+$  and  $Ag^- \leftrightarrow Ag^-$ , whereas *c*-type transitions occur between the tunneling states:  $Ag^+ \leftrightarrow Ag^-$ .



Fig. 4: Fortrat diagram of *n*-propanol. **Top panel:** A broadband measurement in the W-band region is shown, where respective regions with signatures of the Ag conformer, leading to the Fortrat diagram, are highlighted. In some regions, the spectrometer is not as sensitive, e.g. around J = 11, compared to others. Bottom panel: The colored lines in the Fortrat diagram mark linkages of  ${}^{q}R_{K_{a}}$  series transitions,  $J_{K_{a},K_{c}} \leftarrow (J-1)_{K_{a},K_{c}-1}$ . DM-DR spectroscopy secured assignments in the W-band region (from J = 10 to J = 16). The existence of two tunneling states is proven as for each  $K_a$  partner series are found; the tunneling states are marked by  $0^+$  and  $0^-$ . More series can be found, but are not assigned yet, e.g.  $^{q}R_3$  ones. We note the larger tunneling splitting of  $^{q}R_6$  series (green lines), e.g., in comparison to  $^{q}R_5$  ones, and the different slopes for  $^{q}R_8$  series (thick purple lines).

Table 2: Spectroscopic parameters of Aa n-propanol.<sup>a</sup>

Parameter Theory		Theory	Experiment			
(MHz)		Ref. K10	Ref. DS81	This work		
Α		26 579	26401.671(50)	26 401.4395(40)		
В		3784	3802.154(11)	3802.16025(39)		
С		3531	3549.543(20)	3549.45427(33)		
$\Delta_K$	x 10 <sup>3</sup>	58.7	141.5(47)	149.31(55)		
$\Delta_{JK}$	x 10 <sup>3</sup>	-2.35	9.57(99)	-0.3608(295)		
$\Delta_J$	x 10 <sup>3</sup>	0.828	1.672(73)	0.99076(48)		
$\delta_K$	x 10 <sup>3</sup>	0.361	-29.6(69)	-3.250(164)		
$\delta_J$	x 10 <sup>3</sup>	0.0737	0.0943(45)	0.129 324(170)		
$\Phi_K$	x 10 <sup>6</sup>	-	_	-153.3(202)		
$\Phi_{KJ}$	x 10 <sup>6</sup>	-	-	-117.60(210)		
$\Phi_{JK}$	x 10 <sup>6</sup>	-	-	4.870(202)		
$\Phi_J$	x 10 <sup>6</sup>	-	-	-0.00390(48)		
$\phi_K$	x 10 <sup>6</sup>	-	-	707.1(212)		
$\phi_{JK}$	x 10 <sup>6</sup>	-	-	1.850(44)		
$\phi_J$	x 10 <sup>6</sup>	-	-	-0.002675(230)		
$L_{JK}$	x 10 <sup>9</sup>	-	_	-119.9(37)		
$L_{JJK}$	x 10 <sup>9</sup>	-	-	-1.015(39)		
$L_J$	x 10 <sup>9</sup>	-	-	-0.002291(152)		
$l_J$	x 10 <sup>9</sup>	-	-	-0.001 340(85)		
<( <i>i</i> , <i>a</i> )	/°	-	29(1)	26.49(38)		
ρ		-	-	0.149 24(107)		
$\epsilon_{10}$		-	-	-1.2696(127)		
Number of transitions <sup>b</sup>			113	928		
Number of lines			73	689		
rms / kHz			195	48		
wrms <sup>b</sup>			-	1.07		

References: K10 (Kisiel et al. 2010); DS81 (Dreizler & Scappini 1981) <sup>a</sup> Watson's A reduction Hamiltonian of asymmetric rotors is used with ERHAM. *A*, *B*, and *C* are rotational constants;  $\Delta$ 's,  $\Phi$ 's, and *L*'s are quartic, sextic, and octic centrifugal distortion constants, respectively. The angle  $\langle (i, a), \text{ the unitless } \rho \text{ value, and } \epsilon_{10} \text{ describe splittings due to}$ the internal rotation of the  $-CH_3$  group.

<sup>b</sup> We remeasured 34 transitions of Dreizler & Scappini (1981) and unweighted 6 *Q*-branch *a*-type transitions. In total, we assigned 1255 transitions.

 $(4 \le K_a \le 9)$  proving the existence of  $0^+$  and  $0^-$  tunneling states (see Fig. 4). The nomenclature  $0^+$  and  $0^-$  is used to highlight that the vibrational ground state v = 0 is observed. We assigned the tunneling state to transitions with  $4 \le K_a \le 9$  in the way that we assumed  $0^+$  transitions to be always lower in frequency than ones of  $0^-$ , which is usually the case if identical transitions are compared (transitions with same J's,  $K_a$ 's, and  $K_c$ 's).

A further confirmation of unambiguous assignments is the observed asymmetry splitting for  ${}^{q}R_{4}$ ,  ${}^{q}R_{5}$ ,  ${}^{q}R_{6}$ , and  ${}^{q}R_{7}$  series. The two asymmetry components, transitions with identical J and  $K_{a}$  but either  $K_{c} = J - K_{a}$  or  $K_{c} = J - K_{a} + 1$ , are frequently blended for low J and higher  $K_{a}$  values (prolate pairing), but this degeneracy is lifted for higher J (cf. colored lines in Fig. 4 which are split into two lines at higher J). The asymmetry splitting is the frequency difference of the u a-type transitions with  $J_{K_{a},J-K_{a}} \leftarrow J - 1_{K_{a},J-K_{a}-1}$  and  $J_{K_{a},J-K_{a}+1} \leftarrow J - 1_{K_{a},J-K_{a}}$ . The predicted asymmetry splittings, from the ab initio calculations of Kisiel et al. (2010), show the right order of magnitude when compared to the observed ones of both tunneling states; more importantly, they show similar systematic deviations for all  $K_{a}$  numbers (see Fig. 5).

At this point, we tried to reproduce the determined frequencies of assigned transitions with  $4 \le K_a \le 9$  by applying a common asymmetric rotor Hamiltonian, which is comparable to what has been done for the *Aa* conformer (cf. Table 2); This time the *SPFIT* program was used (Pickett 1991). Many series seem to deviate from asymmetric rotor spectra. We expect identical  ${}^{q}R_{K_{a}}$  series (same  $K_{a}$  and same asymmetry side) of the two tunneling states to appear as nearly parallel lines in Fig. 4, whose spacing should decrease for increasing  $K_{a}$  quantum numbers. This is not the case, for instance, the two series with  $K_{a} = 8$  (purple lines) have quite different slopes and the tunneling splitting of series with  $K_{a} = 6$  (green lines) is larger than that of the series with  $K_{a} = 5$  (red lines). We tried to incorporate even order Coriolis-type parameters ( $F_{bc}$ ,  $F_{ac}$ , and  $F_{ab}$ ) to account for possible tunneling-rotation interaction between the two tunneling states,  $Ag^{+}$  and  $Ag^{-}$ , as has been done successfully several times for similar molecular systems in the past; for instance, in Pickett (1972); Christen & Müller (2003); Pearson et al. (1997); Zingsheim et al. (2022). However, for the case of Ag *n*-propanol, still no experimental accuracy for our predictions has been reached.

We continued assigning J,  $K_a$ , and  $K_c$  quantum numbers to transitions with  $K_a < 3$ . However, transitions from one of the two series,  $Ag^+$  and  $Ag^-$ , are not strictly lower in frequency than the other, in contrast to series with  $4 \le K_a \le 9$ . Crossing series are observed in Loomis-Wood plots (Figs. B.2-B.5 in the appendix). Therefore, the tunneling state assignments remain speculative for transitions with  $K_a < 3$ . The two tunneling series ( $^{q}R$ series with identical  $K_a$  and  $K_c$  quantum numbers) with  $K_a < 3$ are even less parallel than observed ones with  $4 \le K_a \le 9$ . We expect states with  $K_a < 3$  to be strongly influenced by tunnelingrotation interaction between  $Ag^+$  and  $Ag^-$  tunneling states, cf. Fig. C.2. Finally, we could not securely assign  $^{q}R_3$  transitions as too many candidate series are close in frequency (Fig. B.6).

In summary, we could prove the existence of the two tunneling states  $Ag^+$  and  $Ag^-$  and assign lines with  $J_{\text{max}} = 69$  and  $K_{a,\text{max}} = 9$ . A quantum number overview of assigned transitions is visualized in Fig. C.1. Deriving a spectroscopic quantum mechanical model with experimental accuracy is beyond the scope of this work because of the multiple perturbations whose origins are unknown at present.

# 4. Discussion

In this work, we present an extended spectroscopic quantum mechanical model based on additional assignments in the MMW region for Aa n-propanol and secured assignments of both tunneling states,  $Ag^+$  and  $Ag^-$ , for Ag n-propanol. In the following, we discuss the spectroscopic results of the Aa and Ag conformers (Sects. 4.1 and 4.2). Subsequently, we offer some ideas to overcome certain limitations, specifically, thanks to a global analysis of both conformers (Sect. 4.3). Next, spectroscopic results of *n*-propanol are judged in regard of astronomical searches (Sect. 4.4). The rotational data are used in a companion article for its astronomical detection (Sect. 4.5). Moreover, we showed that DM-DR measurements can straightforwardly be extended to frequencies higher than the W-band, see Fig. 3, as was demonstrated in the original DM-DR manuscript (Zingsheim et al. 2021).

#### 4.1. Aa n-propanol

We extended spectroscopic assignments of Aa *n*-propanol up to 505 GHz with  $J_{\text{max}} = 70$  and  $K_{a,\text{max}} = 6$ . Finally, 928 transitions are well reproduced by the spectroscopic parameters in Table 2. Larger systematic deviations  $\Delta v$  from our predictions occur for certain series (cf. also Fig. 2); deviations start from transitions with similar J when series with identical  $K_a$  values are compared. The deviations occur at rather low  $K_a$  values and the origins of these perturbations are unknown so far. Transitions



Fig. 5: Asymmetry splitting of the  ${}^{q}R_{4}$ ,  ${}^{q}R_{5}$ ,  ${}^{q}R_{6}$ , and  ${}^{q}R_{7}$  series transitions. (a) Exemplary Loomis-Wood plot for  ${}^{q}R_{6}$  transitions of 0<sup>+</sup>. Two transitions, either with  $K_{c} = J - K_{a} + 1$  or with  $K_{c} = J - K_{a}$ , are observed and marked by green stars in each row. The average frequency of both transitions is the center frequency (Frequency Offset is 0 MHz). Screenshot of the *LLWP* software (Bonah et al. 2022). The frequency difference of both components is the asymmetry splitting for  $K_{a} = 6$  of 0<sup>+</sup> for a given J. (b) Observed and calculated asymmetry splittings for  ${}^{q}R_{4}$  (blue),  ${}^{q}R_{5}$  (red),  ${}^{q}R_{6}$  (green) and  ${}^{q}R_{6}$  (orange). Similar observed and calculated asymmetry splittings confirm the correct assignment of  $K_{a}$  quantum numbers. We note that small asymmetry splittings are not resolvable, both transitions are partly blended, and missing experimental values for J = 17 - 22 originate in the experimental gap from 130–170 GHz.

involving  $K_a = 0$  and 1 seem to be unperturbed as *b*-type transitions involving  $K_a = 1 \leftrightarrow 0$  are well reproduced in our final spectroscopic model. On the other hand, *b*-type transitions with  $K_a = 2 \leftrightarrow 1$ ,  $K_a = 3 \leftrightarrow 2$ , and  $K_a = 4 \leftrightarrow 3$  are only well reproduced up to around J = 30, J = 20, and J = 10, respectively (see Table 1). This case is similar to that of *a*-ethanol, as discussed for the  $^{13}$ C isotopomers (Bouchez et al. 2012).

If different states are interacting, consecutive transitions of one series are frequently perturbed. In particular, systematic deviations occur around the strongest perturbed energy levels. Socalled avoided crossing patterns may be observed; in this case, transitions belonging to the same series with somewhat smaller and larger quantum numbers than heavily perturbed ones can often be assigned with confidence and, more importantly, reproduced to experimental accuracy without treating the interaction. In this way, perturbed energy levels can be located and transitions from these levels can be unweighted in the analysis if the origin of the interaction is unknown or indescribable. For instance, at the start of the analysis of gauche-propanal, certain v = 0 series had to be unweighted (Zingsheim et al. 2017), but they could be properly described by including Fermi resonances and Coriolis interaction in a global analysis with  $v_{24} = 1$  (Zingsheim et al. 2022). In the case of Aa n-propanol, no transitions with higher J values than the ones that deviate from the prediction could be fit to experimental accuracy; the strongest perturbed energy levels and therefore the center of the perturbation, or the strongest perturbed levels and their exact energies, are unknown to this point. Therefore, we expect that our final spectroscopic model is effective since already perturbed transitions may be included in the line list.

# 4.2. Ag n-propanol

We thoroughly examined assignments of Ag *n*-propanol in the MW and MMW regions and proved the existence of the two tunneling states  $Ag^+$  and  $Ag^-$ . First, DM-DR measurements linked various  ${}^{q}R$  series transitions and assignment of these series was undertaken with the support of the *LLWP* software. We could not confirm doublet patterns from the literature (cf. Fig. B.1; Abdurakhmanov et al. 1976), but we did find others instead (see Fig. 4). The ab initio calculations (Kisiel et al. 2010) allowed us to pinpoint patterns and, in particular, the verification of assignments by studying the asymmetry splitting, cf. Fig. 5. In this way, assignments up to  $J_{max} = 69$  and  $K_{a,max} = 9$  with frequencies up to 505 GHz have been achieved. A derivation of a spectroscopic quantum mechanical model to fit the experimental data accurately is beyond the scope of this work.

#### 4.3. Global analysis: Aa, Ag<sup>+</sup>, and Ag<sup>-</sup> conformers

In the future, *Aa n*-propanol should be described in a more sophisticated spectroscopic quantum mechanical model and an initial description of *Ag n*-propanol with experimental accuracy should be derived. This may be possible only through a global analysis of *Aa*, *Ag*<sup>+</sup>, and *Ag*<sup>-</sup> conformers. The rotational energy levels of the *G* family seem to be isolated from *A* conformers ones (Kisiel et al. 2010), therefore, an additional incorporation of *G* conformers in a global analysis is expected to be unnecessary. A global analysis of the *A* family is expected to be successful only if Coriolis interaction and Fermi resonance between the *Aa* conformer and *Ag*<sup>+</sup> or *Ag*<sup>-</sup> tunneling states and, very importantly, also tunneling-rotation interactions between the *Ag*<sup>+</sup> and *Ag*<sup>-</sup> tunneling states are considered.<sup>8</sup> The first interaction terms which should be investigated are the ones used to describe

<sup>&</sup>lt;sup>8</sup> Future analyses may also benefit by taking into account preliminary results of Kisiel et. al., presented at the International Symposium on Molecular Spectroscopy (ISMS) 2006. We were made aware of this after submission; assignments are not publicly available.

the system anti, gauche<sup>+</sup>, and gauche<sup>-</sup> ethanol (Pearson et al. 2008), as the group theoretical considerations and probably also the energy splittings are similar.<sup>9</sup> A quantum mechanical treatment taking interactions into account would also allow to determine relative energies as has been done for the *G* family (Kisiel et al. 2010). Furthermore, the potential of the -OH group rotation may be studied in detail by taking also higher vibrationally excited states into account.

# 4.4. Astronomical search of n-propanol

The partition function used for creating the predictions of all five conformers, given in the supplementary material, is  $Q_{Gx} = 513984.0861$  (*log Q* = 5.7109) at 300 K.<sup>10</sup> This value is derived from the extensive parameter set from Kisiel et al. (2010) and takes the three *G* conformers and their relative energies into account. Thereby, the degeneracy factor was set to 4 (2 for the internal rotation of the –CH<sub>3</sub> group plus 2 when taking the mirror images into account). Conformational correction factors taking into account the existence of the *A* conformers as well as vibrational correction factors, can be found in Table A.1 in the appendix.

Predictions of Ga, Gg, and Gg' are based on the parameter set of Kisiel et al. (2010). We note that the relative energy of Ga was set to  $E_{Ga} = 0 \text{ cm}^{-1}$  (consequently,  $E_{Gg} = 47.8 \text{ cm}^{-1}$ and  $E_{Gg'} = 50.8 \text{ cm}^{-1}$ ), instead of using Gg as reference point, as done in the original data. This adaption is inevitable to derive meaningful intensities and lower state energies in the predictions. Predictions in Pickett's SPCAT format (\*.*cat* files; Pickett 1991) can be found in the supplementary material and will be added to the CDMS as well.

The predictions of *Aa n*-propanol are limited to transitions with *J* up to 70, 30, 20, and 10 for  $K_a = 1, 2, 3$ , and 4, respectively. All assigned transitions with  $K_a \leq 5$  can be used in astronomical observations, but measured transition frequencies should replace predicted ones. Transitions with  $K_a > 5$  should not be used in astronomical observations due to the limitations of our spectroscopic quantum mechanical model. The intensities and lower state energies in our predictions were corrected by an assumed relative energy of  $E_{Aa} = 30 \text{ cm}^{-1}$ . Overall, the number of unambiguously assigned transitions up to 505 GHz and the quantum number coverage for low  $K_a$  values is sufficient to search for rotational transitions of the *Aa* conformer in space.

An astronomical detection of Ag *n*-propanol is facilitated thanks to many unambiguously assigned transitions. For creating meaningful predictions, despite the fact that a quantum mechanical model with experimental accuracy is missing, we used, in a first step, ab initio calculations that are based on Kisiel et al. (2010) to derive theoretical predictions. Even though predicted frequencies do not match experimental accuracy at all, meaningful physical parameters, such as the intensities, lower-state energy, and degeneracies of transitions are obtained by assuming a relative energy of  $E_{Ag} = 45.5$  cm<sup>-1</sup>. Finally, we replaced the predicted frequencies by the measured ones and neglected all unassigned transitions. In doing so, an astronomical detection of Ag *n*-propanol is possible, but deriving a proper spectroscopic quantum mechanical model should be done in particular to derive relative energies and to enlarge the number of usable transitions.

#### 4.5. Astronomical detection of n-propanol

Some members of our team searched for distinctive rotational signatures of all five conformers of *n*-propanol in a companion astronomical study that uses the survey called Re-exploring Molecular Complexity with ALMA (ReMoCa; details given in Belloche et al. 2019) toward Sagittarius (Sgr) B2(N) (Belloche et al. 2022). This astronomical study reports the identification of *n*-propanol toward the secondary hot molecular core, Sgr B2(N2), thanks to the spectroscopic predictions from this work. The detection relies on eight and five lines of the Gg' and Ag conformers, respectively. The catalogs produced in this work were instrumental in securing this first astronomical detection of *n*-propanol in a hot core.

# 5. Conclusion

We recorded spectra of *n*-propanol in the region of 18-505 GHz, which allowed us to analyze the rotational signatures of its *Aa* and *Ag* conformers, in addition to the already extensively studied *Ga*, *Gg*, and *Gg*' conformers (Kisiel et al. 2010). Therefore, rotational signatures of all five conformers of *n*-propanol are available at present and the main conclusions of this work are as follows:

- 1. An extended quantum mechanical model of rotational spectra of *Aa n*-propanol has been made available.
- 2. The existence of the two tunneling states,  $Ag^+$  and  $Ag^-$ , is proven for the first time by unambiguously assigning transitions of Ag *n*-propanol.
- 3. A global analysis of *Aa* and *Ag* conformers is recommended if the spectroscopic description of the *Aa* conformer needs to be extended or if a quantum mechanical model with experimental accuracy of the *Ag* conformer is the target.
- 4. Predictions of the rotational spectra of all five conformers of *n*-propanol (\*.*cat* files in the supplementary material) enable their astronomical search.

In addition, the spectroscopic analysis presented in this work allowed for the astronomical detection of *n*-propanol toward a hot core in a companion article (Belloche et al. 2022).

Acknowledgements. We would like to thank Dr. Matthias Justen and Dr. Leonid Surin for their help in translating Russian manuscripts and providing literature data from the VINITI database. B. H. acknowledges financial support by the Deutsche Forschungsgemeinschaft (DFG; project ID WE 5874/1-1). This work has been supported via Collaborative Research Centre 956, sub-projects B3 and B4, and the "Cologne Center for Terahertz Spectroscopy", both funded by the DFG (project IDs 184018867 and SCHL 341/15-1, respectively).

#### References

Abdurakhmanov, A., Ragimova, R., Veliyulin, E., & Imanov, L. 1976, VINITI (a russian database), 3883-76, 6pp Abdurakhmanov, A., Rahimova, R., & Imanov, L. 1968, Izv. Akad. Nauk. Azerb.

Abdurakhmanov, A., Rahimova, R., & Imanov, L. 1968, Izv. Akad. Nauk. Azerb. SSR, 3, 15

- Abdurakhmanov, A., Rahimova, R., & Imanov, L. 1970, Phys. Lett. A, 32, 123
   Ball, J. A., Gottlieb, C. A., Lilley, A. E., & Radford, H. E. 1970, Astrophys. J. Lett., 162, L203
- Belloche, A., Garrod, R. T., Müller, H. S. P., et al. 2019, A&A, 628, A10
- Belloche, A., Garrod, R. T., Zingsheim, O., Müller, H. S. P., & Menten, K. M. 2022, A&A, submitted
- Bonah, L., Zingsheim, O., Müller, H. S., Guillemin, J.-C., & Schlemmer, S. 2022, J. Mol. Spectrosc., to be submitted soon, download the *LLWP* program at https://llwp.astro.uni-koeln.de/

<sup>&</sup>lt;sup>9</sup> Caution is advised concerning the number of interaction parameters; Using too many leads to high correlations and convergence problems. Using too few may lead to a more stable fit, but may not account for all perturbations, or, if it does, at the expense of unphysical parameters. <sup>10</sup> Our predictions (\*.*cat* files) will be available together with further

<sup>&</sup>lt;sup>10</sup> Our predictions (\*.*cat* files) will be available together with further information; cf. Appendix A, in the CDMS soon.

- Bouchez, A., Walters, A., Müller, H. S., et al. 2012, J Quant Spectrosc Radiat Transf, 113, 1148
- Christen, D. & Müller, H. S. P. 2003, Phys. Chem. Chem. Phys., 5, 3600
- Dreizler, H. & Scappini, F. 1981, Z. Naturforsch. A, 36, 1187
- Endres, C. P., Schlemmer, S., Schilke, P., Stutzki, J., & Müller, H. S. 2016, J.
- Mol. Spectrosc., 327, 95, New Visions of Spectroscopic Databases, Volume II
- Fukushima, K. & Zwolinski, B. J. 1968, J. Mol. Spectrosc., 26, 368
- Groner, P. 1992, J. Mol. Spectrosc., 156, 164
- Groner, P. 1997, J. Chem. Phys., 107, 4483
- Groner, P. 2012, J. Mol. Spectrosc., 278, 52
- Hermanns, M., Wehres, N., Lewen, F., Müller, H., & Schlemmer, S. 2019, J. Mol. Spectrosc., 358, 25
- Herschbach, D. R. & Swalen, J. D. 1958, J. Chem. Phys., 29 (4), 761
- Hollis, J. M., Lovas, F. J., Jewell, P. R., & Coudert, L. H. 2002, Astrophys. J., 571, L59
- Johnson, D. R., Lovas, F. J., Gottlieb, C. A., et al. 1977, Astrophys. J. Lett., 218, 370
- Jørgensen, J. K., van der Wiel, M. H. D., Coutens, A., et al. 2016, A&A, 595, A117
- Kahn, K. & Bruice, T. C. 2005, ChemPhysChem, 6, 487
- Kisiel, Z., Dorosh, O., Maeda, A., et al. 2010, Phys. Chem. Chem. Phys., 12, 8329
- Lotta, T., Murto, J., Räsänen, M., & Aspiala, A. 1984, Chem. Phys., 86, 105
- Maeda, A., Lucia, F. C. D., Herbst, E., et al. 2006, Astrophys. J., Suppl. Ser., 162, 428
- Manigand, S., Jørgensen, J. K., Calcutt, H., et al. 2020, A&A, 635, A48
- Martin-Drumel, M.-A., van Wijngaarden, J., Zingsheim, O., et al. 2015, J. Mol. Spectrosc., 307, 33
- Müller, H. S., Schlöder, F., Stutzki, J., & Winnewisser, G. 2005, J. Mol. Struct., 742, 215
- Müller, H. S., Thorwirth, S., Roth, D., & Winnewisser, G. 2001, A&A, 370, L49 Ordu, M. H., Zingsheim, O., Belloche, A., et al. 2019, Astron. Astrophys., 629,
- A72 Pearson, J. C., Brauer, C. S., & Drouin, B. J. 2008, J. Mol. Spectrosc., 251, 394
- Pearson, J. C., Sastry, K. V. L. N., Herbst, E., & Lucia, F. C. D. 1997, Astrophys.
- J., 480, 420
- Pickett, H. M. 1972, J. Chem. Phys., 56, 1715
- Pickett, H. M. 1991, J. Mol. Spectrosc., 148, 371
- Pickett, H. M. & Scroggin, D. G. 1974, J. Chem. Phys., 61, 3954
- Rubin, R. H., Swenson, G. W., J., Benson, R. C., Tigelaar, H. L., & Flygare, W. H. 1971, Astrophys. J. Lett., 169, L39
- Tercero, B., Cernicharo, J., López, A., et al. 2015, A&A, 582, L1
- Tercero, B., Cuadrado, S., López, A., et al. 2018, A&A, 620, L6
- Zingsheim, O., Bonah, L., Lewen, F., et al. 2021, J. Mol. Spectrosc., 381, 111519
- Zingsheim, O., Müller, H. S. P., Lewen, F., Jørgensen, J. K., & Schlemmer, S. 2017, J. Mol. Spectrosc., 342, 125
- Zingsheim, O., Müller, H. S., Bonah, L., et al. 2022, J. Mol. Spectrosc., 384, 111565
- Zuckerman, B., Turner, B., Johnson, D., et al. 1975, Astrophys. J., 196, L99

# Appendix A: Partition function *Q* of *n*-propanol

We calculated the partition function of the three G conformers  $Q_{\rm Gx}$  with SPCAT to be 513984.0861 at 300 K, which takes into account the Ga, Gg, and Gg' conformers and a degeneracy factor of 4 (a factor of 2 for the internal rotation of the -CH<sub>3</sub> group and another factor of 2 for mirror images of the respective conformers; see next paragraph for more details).<sup>11</sup> This value is based on the quantum mechanical description from Kisiel et al. (2010). The spectroscopic quantum mechanical models of Aa and Agconformers are not as accurate as the ones of the G family and the relative energies of  $E_{Aa}$  and  $E_{Ag}$  to  $E_{Ga}$  are somewhat uncertain, thus, we used the partition function derived from the Gfamily  $Q_{Gx}$  for creating predictions (\*.*cat* files) of all five conformers. For the predictions of Aa and Ag conformers, the degeneracy factor is 2 and 4, respectively. Both conformers have an internal rotor (A and E transition, whereby both components are blended for the Ag conformer), but the Aa conformer does not occur in a mirror configuration as the other four conformers do. In this way, we derived correct relative intensities of all five conformers.

For predicting correct absolute intensities, which are fundamental for deriving meaningful column densities of molecules in space, the existence of both A conformers and all vibrational states need to be taken into account in order to derive  $Q_{\text{Tot}}$ . Therefore, we present conformational and vibrational correction factors in Table A.1. The conformational correction factors are derived as follows:

$$F_{\text{Conf}}^{T} = \frac{Q_{\text{Ga}(1)}^{T} \cdot f_{\text{Aa}}^{'} \cdot f_{\text{Aa}}^{''} \cdot BF_{\text{Aa}}^{T} + Q_{\text{Ga}(1)}^{T} \cdot f_{\text{Ag}}^{'} \cdot f_{\text{Ag}}^{''} \cdot BF_{\text{Ag}}^{T} + Q_{\text{Ga}}^{T}}{Q_{\text{Gx}}^{T}} \\ = \frac{Q_{\text{Ga}(1)}^{T} \cdot 2 \cdot 1 \cdot BF_{\text{Ag}}^{T} + Q_{\text{Ga}(1)}^{T} \cdot 2 \cdot 2 \cdot BF_{\text{Ag}}^{T} + Q_{\text{Gx}}^{T}}{Q_{\text{Gx}}^{T}},$$
(A.1)

where the partition function  $Q_{Ga(1)}^T$  is calculated with degeneracy factor 1, for instance, to be  $Q_{Ga(1)} = 49318.3481$  at 300 K. The first factor f' = 2 takes the internal rotation into account and the second factor, namely, f'' = 1 or 2, the conformer multiplicity of Aa and Ag, respectively. The Boltzmann factor  $BF_{conf}^T = exp(-E_{conf}/k_BT)$  of both conformers is derived by assuming  $E_{Aa} = 30.0 \text{ cm}^{-1}$  and  $E_{Ag} = 45.5 \text{ cm}^{-1}$ . The vibrational correction factors were calculated using fundamental vibrational frequencies (with  $\nu \le 1200 \text{ cm}^{-1}$ ) of the Aa conformer (Fukushima & Zwolinski 1968), that is,  $v_{30} = 131 \text{ cm}^{-1}$ ,  $v_{29} = 240 \text{ cm}^{-1}$ ,  $v_{28} = 286 \text{ cm}^{-1}$ ,  $v_{27} = 332 \text{ cm}^{-1}$ ,  $v_{26} =$  $463 \text{ cm}^{-1}$ ,  $v_{25} = 758 \text{ cm}^{-1}$ ,  $v_{24} = 858 \text{ cm}^{-1}$ ,  $v_{23} = 898 \text{ cm}^{-1}$ ,  $v_{22} = 971 \text{ cm}^{-1}$ ,  $v_{21} = 1013 \text{ cm}^{-1}$ ,  $v_{20} = 1047 \text{ cm}^{-1}$ , and  $v_{19} = 1066 \text{ cm}^{-1}$ . Applying the correction factors from Table A.1 allows to derive accurate column densities of *n*-propanol and derive relative abundances to other molecular species in astronomical sources.

# Appendix B: Rotational signatures of Ag n-propanol

First, a part of the MW spectrum of *n*-propanol and propanal, depicting the  $3_{K_a,K_c} \leftarrow 2_{K_a,K_c-1} a$ -type transitions, recorded with

Article number, page 10 of 18

the CP-FTMW spectrometer, are shown in Fig. B.1. Rotational signatures are shown in the frequency regions 21.5-22.5 GHz and 25.0-26.0 GHz, respectively. Both measurements were performed under similar conditions. The rotational spectrum of gauche-propanal is well-understood (Zingsheim et al. 2022), in particular doublet structures due to two tunnelings states are observed; Two  $K_a = 1$  transitions which are close in frequency and have similar intensities. We note that tunneling-rotation interaction mixes the wave functions of involved rotational energy levels, in this case intensity borrowing may play a role and nominally forbidden transitions can be observed (cf. transitions marked with  $v = 0^+ \leftrightarrow 0^-$ ). The interpretation of the rotational spectrum of Ag n-propanol is more complicated. Assigned doublet transitions for Ag n-propanol from Abdurakhmanov et al. (1976) are marked in red, but cannot be verified in our CP-FTMW jet experiment. However, two transitions with similar intensities close in frequency are found twice. Comparison of the rotational spectra of both molecules leads to tentative assignments of  $K_a = 1$  transitions of Ag *n*-propanol. The appearance of two transitions close in frequency were first hints of the existence of two tunneling states  $(Ag^+ \text{ and } Ag^-)$ , similar to what is observed for gauche-propanal. The different tunneling splittings of transitions with  $K_a = 1$  for Ag n-propanol suggest that involved rotational energy levels are perturbed due to so far untreated interactions. We note that the transition marked with an asterisk (\*) is another possible candidate for a transition with  $K_a = 1$ , if <sup>q</sup>R series are tracked, cf. Fig. B.4. Unambiguously assigning these transitions can be done if a spectroscopic quantum mechanical model with experimental accuracy is derived.

Second, typical spectroscopic signatures of  ${}^{q}R_{K_{a}}$  series transitions allow for unambiguous assignments of the Ag conformer, despite a missing quantum mechanical description. A broader Fortrat diagram than shown in Fig. 4 is presented in Fig. B.2. Then, Loomis-Wood plots are shown for series with  $0 \le K_{a} \le 3$  to further document these assignments (see Figs. B.3–B.6). Series with  $4 \le K_{a} \le 9$ , which have been assigned in this work as well, are already well represented by Fig. 4 in the main article. The Loomis-Wood plots are all screenshots of the LLWP software used for assigning transitions within this work (Bonah et al. 2022).

# Appendix C: Overview of the rotational energy levels of Aa and Ag n-propanol

We present a quantum number overview of our analysis for Aa and Ag conformers of *n*-propanol here as we could not reproduce all assigned transitions to experimental uncertainty. However, all assigned transitions can be used for astronomical searches of the rotational spectra of both conformers, therefore, quantum number overviews are given for clarification (see Fig. C.1). Furthermore, an estimated reduced energy diagram of Ag *n*-propanol is shown, which illustrates the location of tunneling-rotation interactions (see Fig. C.2). Of course, interactions between rotational energy levels of Ag *n*-propanol and those of Aa *n*-propanol should also be considered if a global analysis is performed, but these interactions are not illustrated here.

<sup>&</sup>lt;sup>11</sup> Reducing the total degeneracy factor of 4 is not possible if only one column density for all five conformers of *n*-propanol are to be derived, e.g., in local thermal equilibrium conditions in space, as the Aa conformer does not exist in a mirror image form and A and E internal rotation components are resolved.

Table A.1: Partition function  $Q_{Gx}$  of all three *G* conformers (with spin weight 4),  $Q_{Ga(1)}$  of *Ga* (with spin weight 1), together with conformational correction factors  $F_{Conf}$ , vibrational correction factors  $F_{Vib}$ , and their product  $F_{Tot}$  for various temperatures of interest for an astronomical detection.

T	$Q_{ m Gx}$	$\log Q_{\rm Gx}$	$Q_{\mathrm{Ga}(1)}$	$F_{Conf}$	$F_{Vib}$	$F_{Tot}$
(K)				-		
300	513 984.0861	5.7109	49318.3481	1.4748	6.4531	9.5167
225	318 241.7669	5.5028	32 031.6352	1.4671	3.2754	4.8055
200	260 495.0562	5.4158	26 839.0622	1.4631	2.6370	3.8583
180	217 323.4156	5.3371	22911.7137	1.4590	2.2291	3.2523
160	177 034.0151	5.2481	19 197.9679	1.4537	1.8966	2.7571
150	158 035.3077	5.1988	17 425.1609	1.4504	1.7549	2.5454
140	139 840.8805	5.1456	15710.7948	1.4466	1.6280	2.3551
120	106 007.9447	5.0253	12465.7151	1.4367	1.4146	2.0325
100	75 871.4993	4.8801	9481.9699	1.4221	1.2499	1.7775
75	44 088.0429	4.6443	6158.4144	1.3905	1.1058	1.5376
37.5	11 387.5229	4.0564	2178.3969	1.2546	1.0067	1.2630
18.75	3228.9908	3.5091	771.4138	1.0769	1.0000	1.0770
9.375	1095.9636	3.0398	273.6960	1.0059	1.0000	1.0059



Fig. B.1: Microwave spectra of *n*-propanol and propanal. Assignments of *a*-type transitions with  $3_{K_a,K_c} \leftarrow 2_{K_a,K_c-1}$  are shown for both molecules ( $K_a$ 's are given in the figure). **Top panel:** Rotational spectrum of gauche-propanal with assignments from Zingsheim et al. (2022) marked in blue. **Bottom panel:** Assigned doublet transitions for Ag *n*-propanol from Abdurakhmanov et al. (1976) are marked in red, but cannot be verified in our CP-FTMW jet experiment. Our tentative assignments of  $K_a = 1$  transitions of Ag *n*-propanol are marked in blue. We note that the transition marked with an asterisk (\*) is another possible candidate for a transition with  $K_a = 1$ , more information can be found in the text.



Fig. B.2: Fortrat diagram of *n*-propanol. The colored lines in the Fortrat diagram mark linkages of  ${}^{q}R_{K_{a}}$  series transitions of Ag *n*-propanol,  $J_{K_{a},K_{c}} \leftarrow (J-1)_{K_{a},K_{c}-1}$ . DM-DR spectroscopy secured assignments in the W-band region (from J = 10 to J = 16). Tentative assignments of  $K_{a} = 1$  transitions derived from the CP-FTMW measurement (Fig. B.1) are marked by the blue arrows. Considering found  ${}^{q}R_{1}$  series, another candidate transition is marked by the blue asterisk. A quantum mechanical model with experimental accuracy can solve the ambiguity of these tentative assignments. The gray area depicts the region of the Fortrat diagram in Fig. 4.



Fig. B.3: Loomis-Wood plot around the assigned  ${}^{q}R_{0}$  series transitions in the W-band region of the Ag conformer, which are shown by red sticks.





Fig. B.4: Loomis-Wood plot around the assigned  ${}^{q}R_{1}$  series transitions in the W-band region of the Ag conformer, which are shown by red sticks. The blue asterisk (\*) marks the same transition as is marked in Fig. B.1.



O. Zingsheim et al.: Rotational spectroscopy of n-propanol: Aa and Ag conformers

Fig. B.5: Loomis-Wood plot around the assigned  ${}^{q}R_{2}$  series transitions in the W-band region of the Ag conformer, which are shown by red sticks. The series higher in frequency with  $K_{a} + K_{c} = J + 1$  is probably strongly interacting with a so far unknown state around J = 12.



Fig. B.6: Loomis-Wood plot around assumed  ${}^{q}R_{3}$  series transitions in the W-band region. No assignments with  $K_{a} = 3$  are secured yet, but the existence of such can be seen by the many unassigned transitions in the shown region. Four series close in frequency are expected; Two asymmetry sides for both tunneling states.





Fig. C.1: Quantum number overview of the analysis of (a) *Aa n*-propanol and (b) *Ag n*-propanol. All rotational energy levels of assigned transitions are visualized with the *J* quantum numbers on the *x*-axis,  $K_a$  quantum numbers on the *y*-axis, and  $K_c$  quantum numbers are either depicted by squares or circles for  $K_c = J - K_a$  or  $K_c = J - K_a + 1$ , respectively. Filled markers depict quantum numbers of assigned transitions whose frequencies are fit to experimental accuracy using our final spectroscopic model in Table 2. The empty circles mark assigned transitions which are not included in the final analysis, but can be used for astronomical searches. For *Aa n*-propanol, we note that we do not differentiate between A and E components of each rotational energy level as in overwhelming majority of cases both components are assigned per transition. For *Ag n*-propanol, the rotational energy levels of the 0<sup>+</sup> and 0<sup>-</sup> tunneling states are given in blue and red, respectively. Triangles are used for transitions whose tunneling state assignments have only been guessed so far.



(b) Reduced energy diagram for rotational energy levels with  $K_a \leq 3$ .

20

30

10

Fig. C.2: Reduced energy diagram of Ag *n*-propanol. The reduced energy is  $E_{red} = E_T + E_{rot} - J \cdot (J + 1) \cdot (B + C)/2$ . Thereby, the energies of the rotational levels  $E_{rot}$  are calculated based on ab initio calculations of Kisiel et al. (2010). We used the same parameters for both tunneling states as ab initio calculations are presented for a single state only. We note that actual parameters of the two tunneling states will differ. We set the energy difference of the two tunneling states to  $3 \text{ cm}^{-1}$ , i.e.  $E_T^{0^+} = 0 \text{ cm}^{-1}$  and  $E_T^{0^-} = 3 \text{ cm}^{-1}$ , on the assumption that the splitting is similar to what has been observed for *gauche*-ethanol (Pearson et al. 1997). (a) The reduced energy diagram for rotational energy levels with  $K_a < 9$  shows that series with  $K_a > 5$  may not cross each other. We note that it may happen if different rotational parameters for the two tunneling states are taken into account. (b) The reduced energy diagram for rotational energy levels with  $K_a \le 4$  is a zoom of the *y*-axis of (a). We note that many series cross each other and even though the exact energy difference between the two tunneling states  $Ag^+$  and  $Ag^-$  is unknown so far, many tunneling-rotation interactions occur in this region and determining an accurate energy difference is essential to properly take them into account.

-4

0